

The background features a dark blue gradient with a starry pattern. Overlaid on this are several white circular elements: a large scale on the left with markings from 140 to 260, and several smaller circles with arrows indicating clockwise or counter-clockwise rotation. The main title is centered in large, white, sans-serif font.

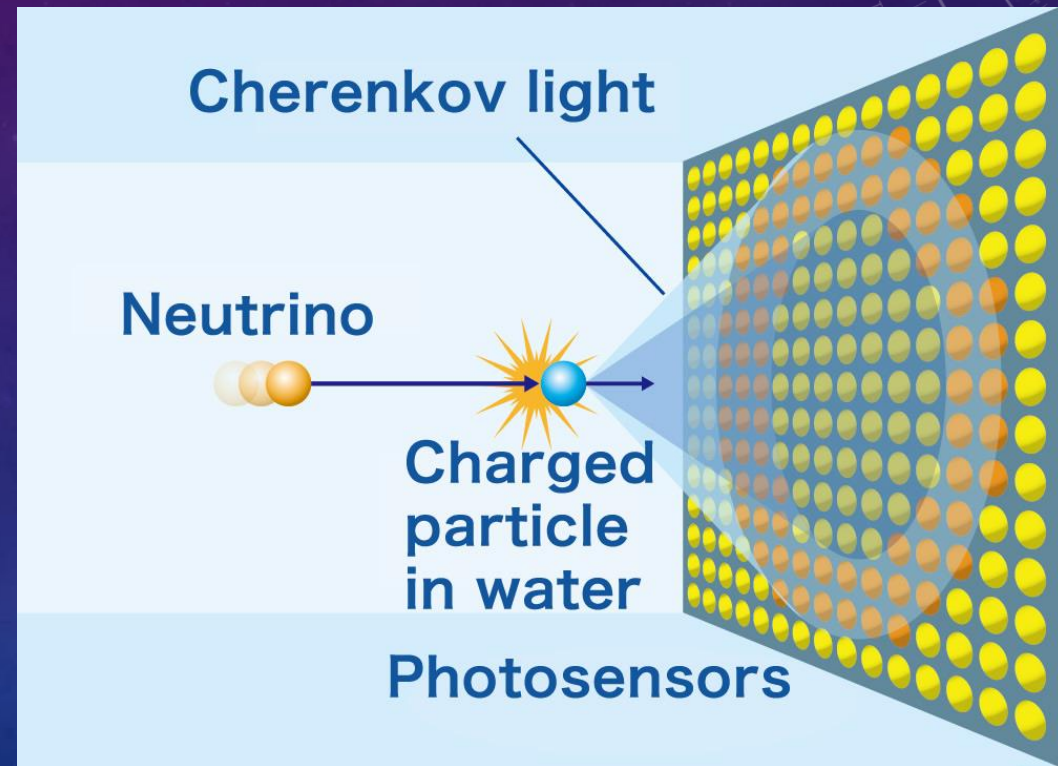
DISCOVERY OF NEUTRINO OSCILLATIONS

FRANKLIN ADAMS

UNIVERSITY OF SOUTH CAROLINA – DEPARTMENT OF PHYSICS AND ASTRONOMY

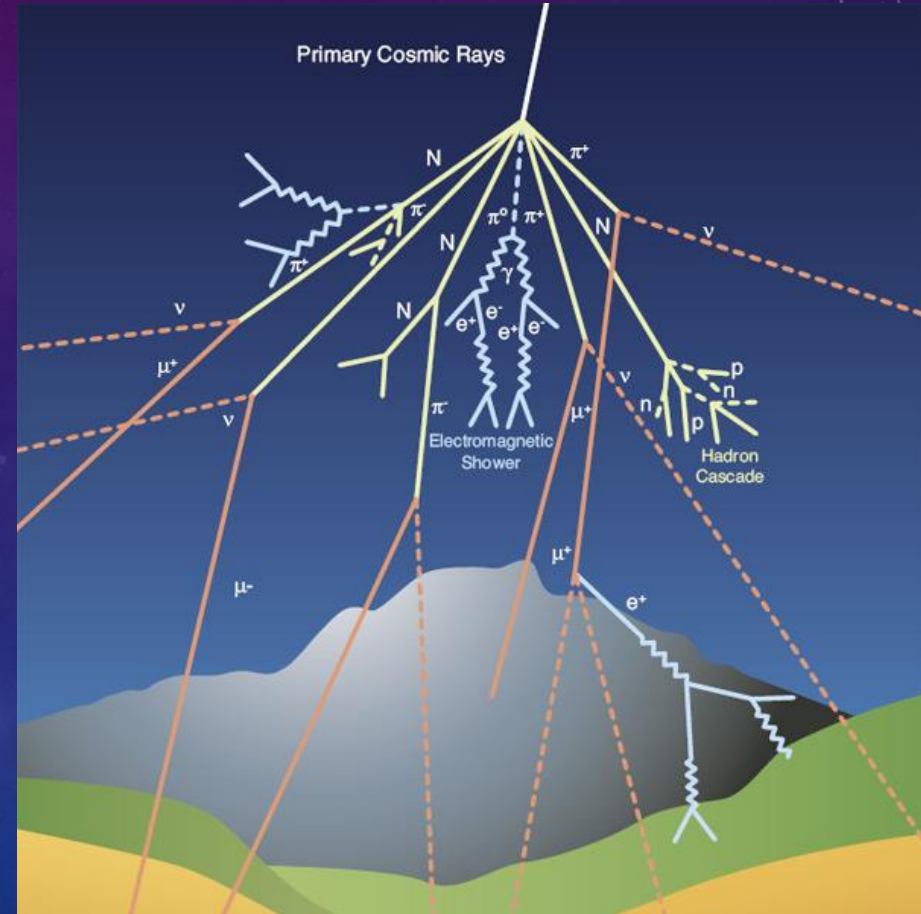
KAMIOKANDE

- Kamioka Neutron Decay Experiment
 - 3000 tons of water
 - 1000 50cm PMTs
 - 1000m underground
- Built in 1983 to detect proton decay, no evidence after 3 years
- Suspected that software was not good enough
- New software was tested on single Cherenkov-ring events
- Unexpected result
 - Much fewer ν_{μ} than expected



COSMIC RAY DECAY

- Cosmic Ray enters the atmosphere
- Produces a $\pi \rightarrow \mu + \nu_\mu$
- $\mu \rightarrow e + \nu_e + \nu_\mu$
- Should expect $\sim 2 \nu_\mu$ per ν_e for atmospheric neutrinos



RESULTS OF KAMIOKANDE

- Monte Carlo simulation 21.8 kiloton*year
- Data from 2.86 kiloton*year exposure
- Not at 2:1 ratio due to detection efficiency of μ detection
- ν_e ratio 1.05 ± 0.11
- ν_μ ratio 0.59 ± 0.07

K. Hirata et al, Phys.Lett.B 205 (1988) 416.

	Data	Prediction
ν_e events	93	88.5
ν_μ events	85	144.0

WHY DOES NEUTRINO OSCILLATION IMPLY MASS

From the time - dependent Schrödinger equation :

$$\begin{aligned} \begin{pmatrix} \nu_1(\vec{x}, t) \\ \nu_2(\vec{x}, t) \end{pmatrix} &= e^{i\vec{p}\cdot\vec{x}} \begin{pmatrix} e^{-iE_1 t} |\nu_1(\mathbf{0})\rangle \\ e^{-iE_2 t} |\nu_2(\mathbf{0})\rangle \end{pmatrix} \\ &= e^{i\vec{p}\cdot\vec{x}} \begin{pmatrix} e^{-iE_1 t} & 0 \\ 0 & e^{-iE_2 t} \end{pmatrix} \begin{pmatrix} |\nu_1(\mathbf{0})\rangle \\ |\nu_2(\mathbf{0})\rangle \end{pmatrix} \end{aligned}$$

Using the relation between mass and flavor eigenstates :

$$\begin{pmatrix} |\nu_\mu(\vec{x}, t)\rangle \\ |\nu_\tau(\vec{x}, t)\rangle \end{pmatrix} = e^{i\vec{p}\cdot\vec{x}} \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} e^{-iE_1 t} & 0 \\ 0 & e^{-iE_2 t} \end{pmatrix} \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} |\nu_\mu(\mathbf{0})\rangle \\ |\nu_\tau(\mathbf{0})\rangle \end{pmatrix}$$

If $|\nu_\mu(\mathbf{0})\rangle = 1$ and $|\nu_\tau(\mathbf{0})\rangle = 0$:

$$||\nu_\tau(\vec{x}, t)\rangle|^2 = \sin^2(2\theta) \sin^2 \frac{(E_2 - E_1)t}{2} \equiv P(\nu_\mu \rightarrow \nu_\tau)$$

WHY DOES NEUTRINO OSCILLATION IMPLY MASS

If $E_1, E_2 \gg m_1, m_2$:

$$E_2 - E_1 = \sqrt{m_2^2 + p^2} - \sqrt{m_1^2 + p^2} \approx \frac{m_2^2 - m_1^2}{2p}$$

and

$$\begin{aligned} t &\approx |\vec{x}| \equiv L, \\ p &\approx E \end{aligned}$$

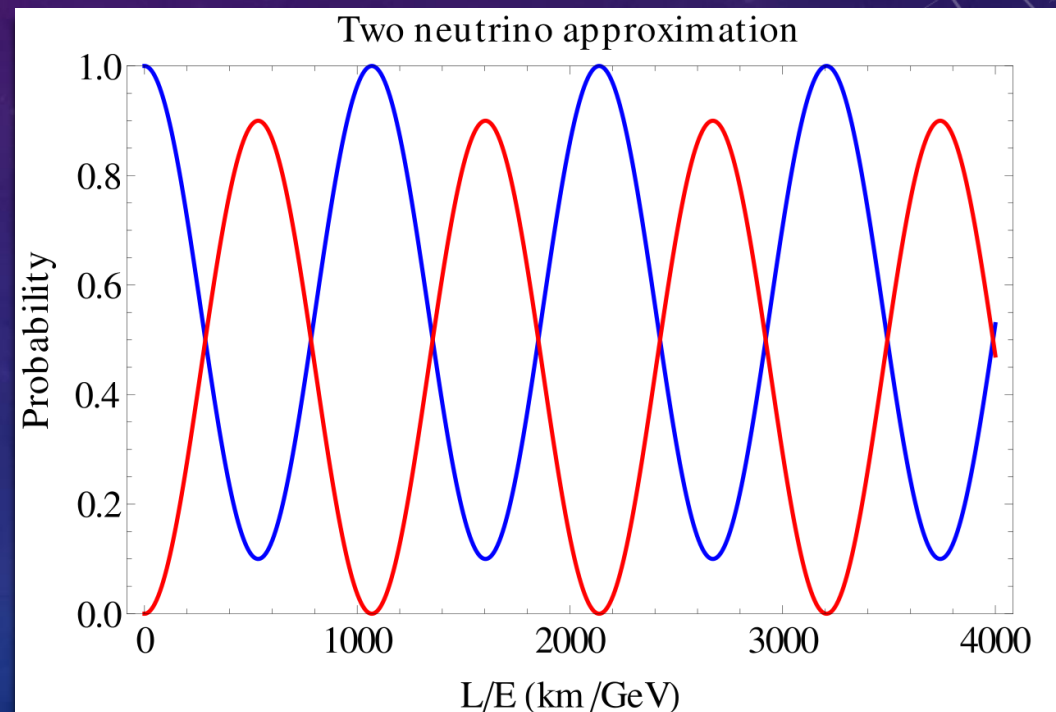
Therefore:

$$P(\nu_\mu \rightarrow \nu_\tau) \approx \sin^2(2\theta) \sin^2\left(\frac{\Delta m^2 L}{4E}\right)$$

- So, the probability of a neutrino oscillating is a function of: θ , L , E , and Δm
- If $\Delta m = 0$ neutrinos will have a zero probability to oscillate

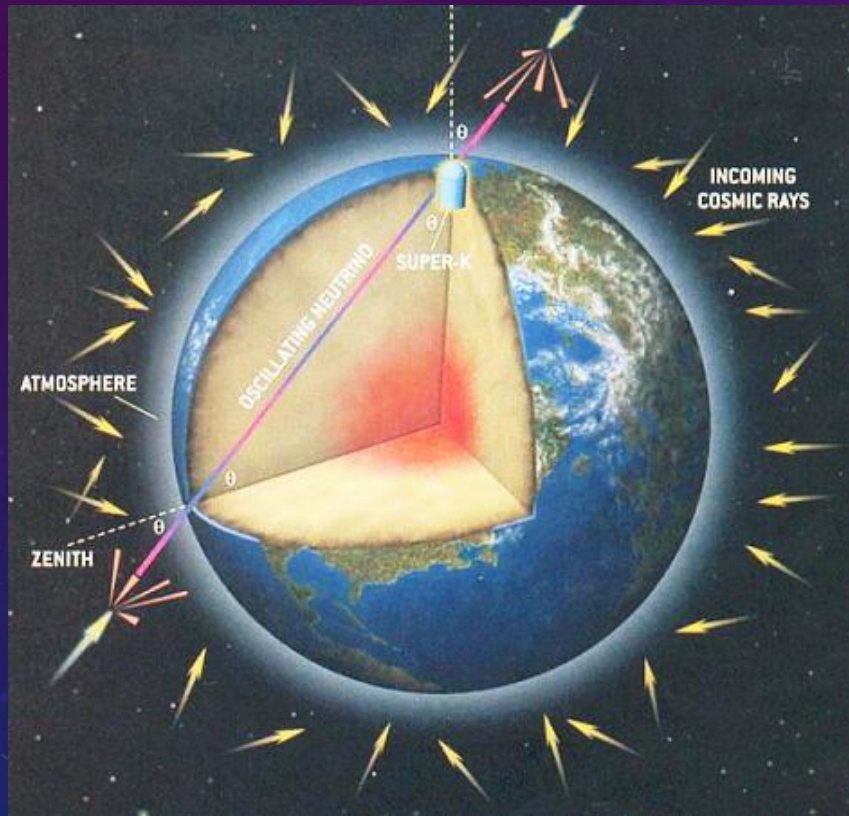
NEUTRINO OSCILLATION CONDITIONS

- In the case of $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillation
 - Blue curve represents probability to remain ν_{μ}
 - Red curve represents probability to become ν_{τ}
- At short distances L , neutrinos will have a very low probability to oscillate



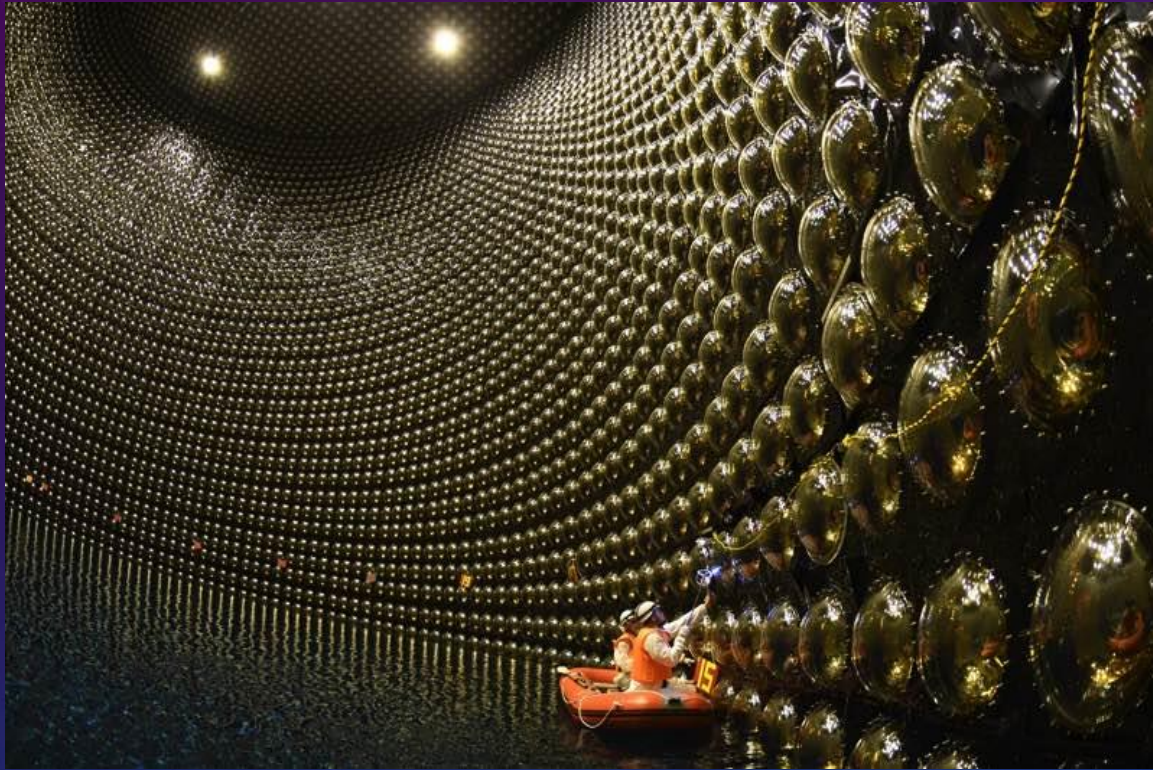
© Wikipedia, Neutrino oscillation 2011

NEUTRINO OSCILLATION CONDITIONS



- Should observe a deficit in ν_{μ} that pass through the earth
 - Upward going ν_{μ}
- Want to create conditions to observe oscillations
- Kamiokande was not enough to be conclusive
- Need a much larger detector

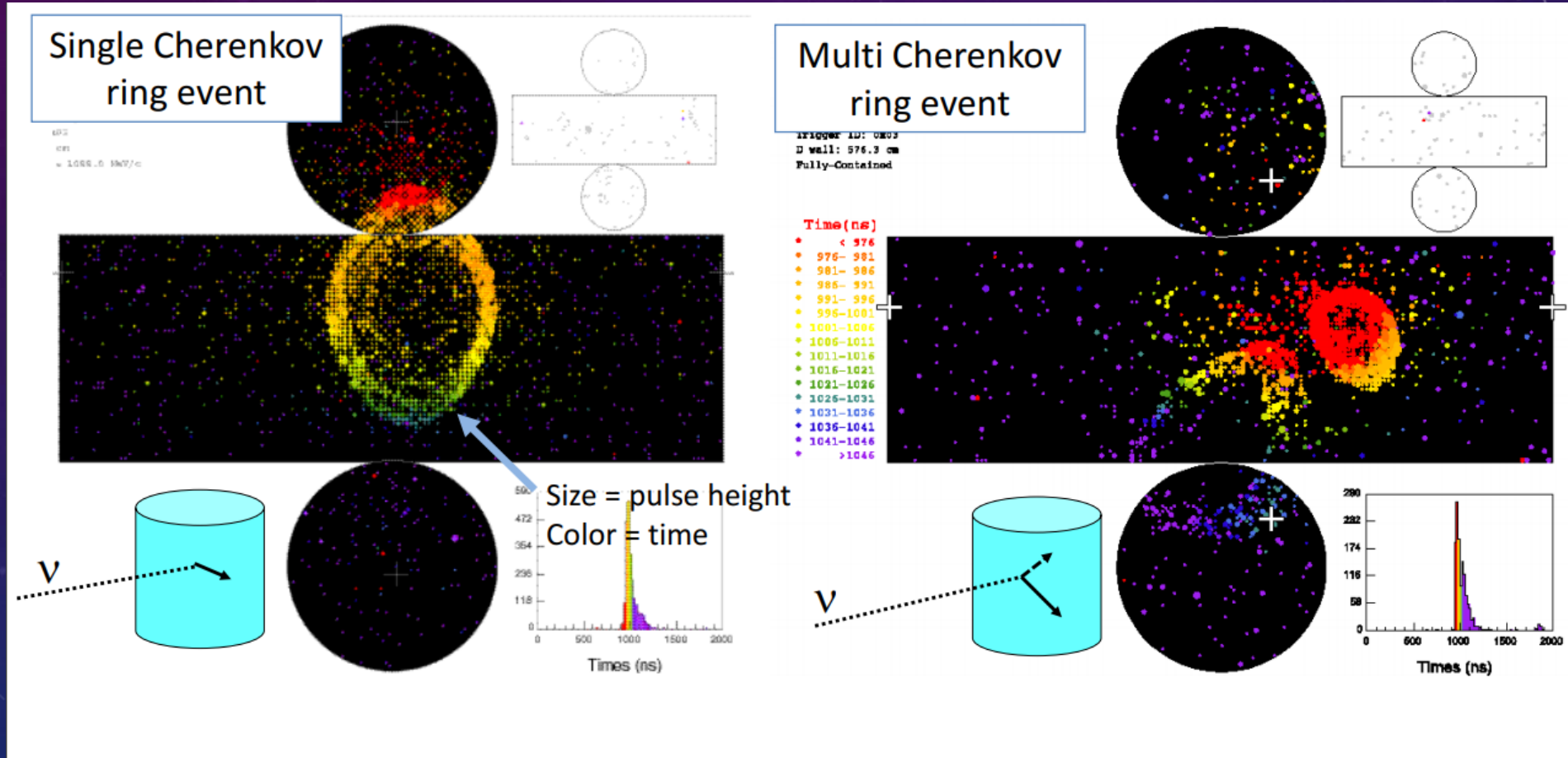
SUPER-KAMIOKANDE



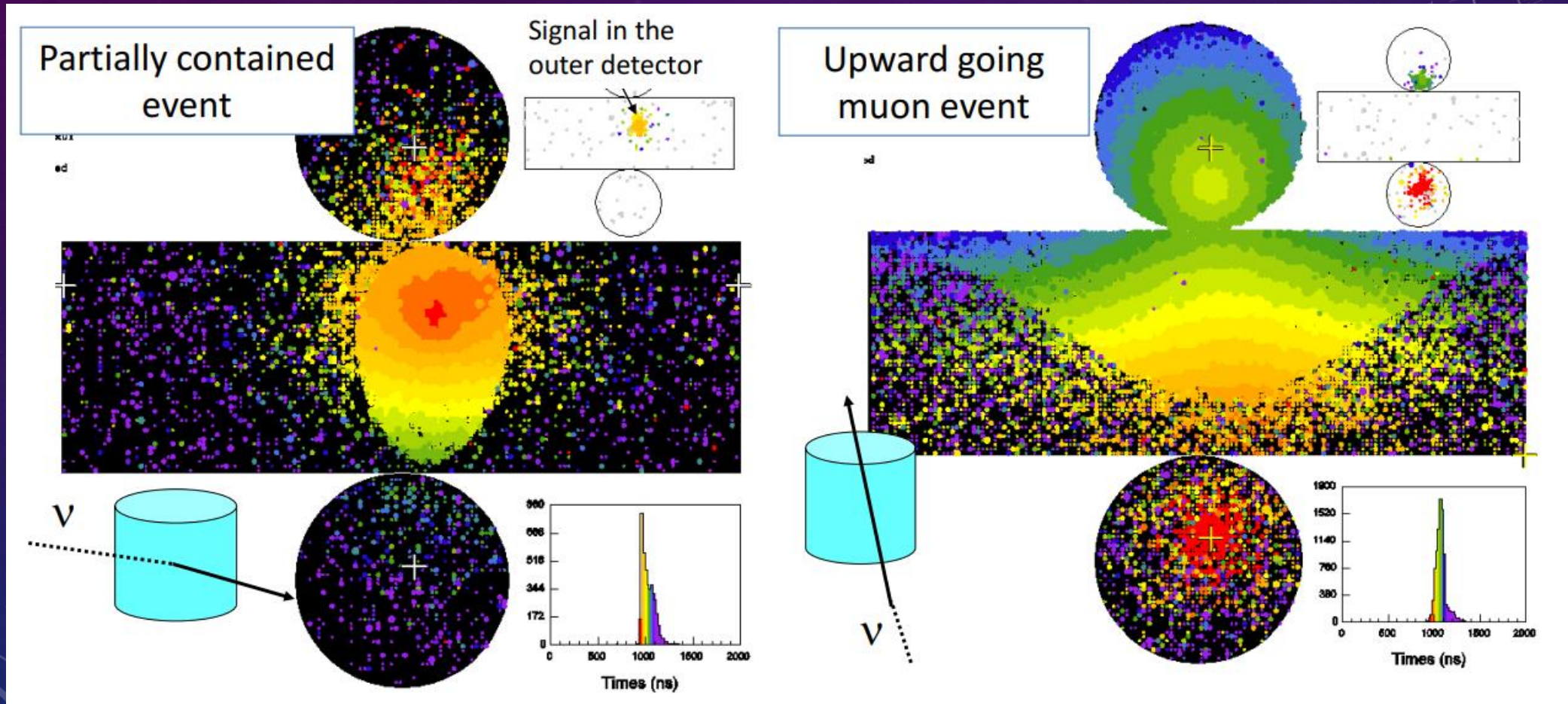
© Kamioka Observatory 2018

- 3000 -> 50,000 ton water Cherenkov Detector
- 1000 -> 13,000 PMT
- 1000m underground

POSSIBLE NEUTRINO EVENTS



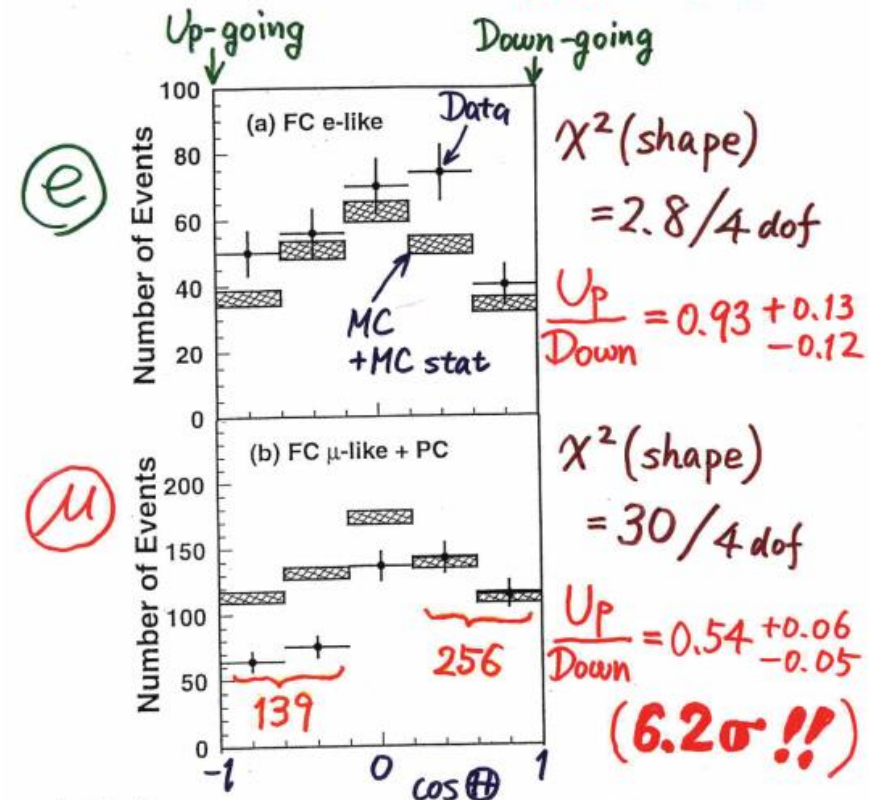
POSSIBLE NEUTRINO EVENTS



RESULTS

- $\cos\theta = 1$ down-going neutrino
- Shaded boxes are predictions
- Crosses are observations
- All things considered 6.2σ
 - ~ 1 in 1.8 billion

Zenith angle dependence (Multi-GeV)



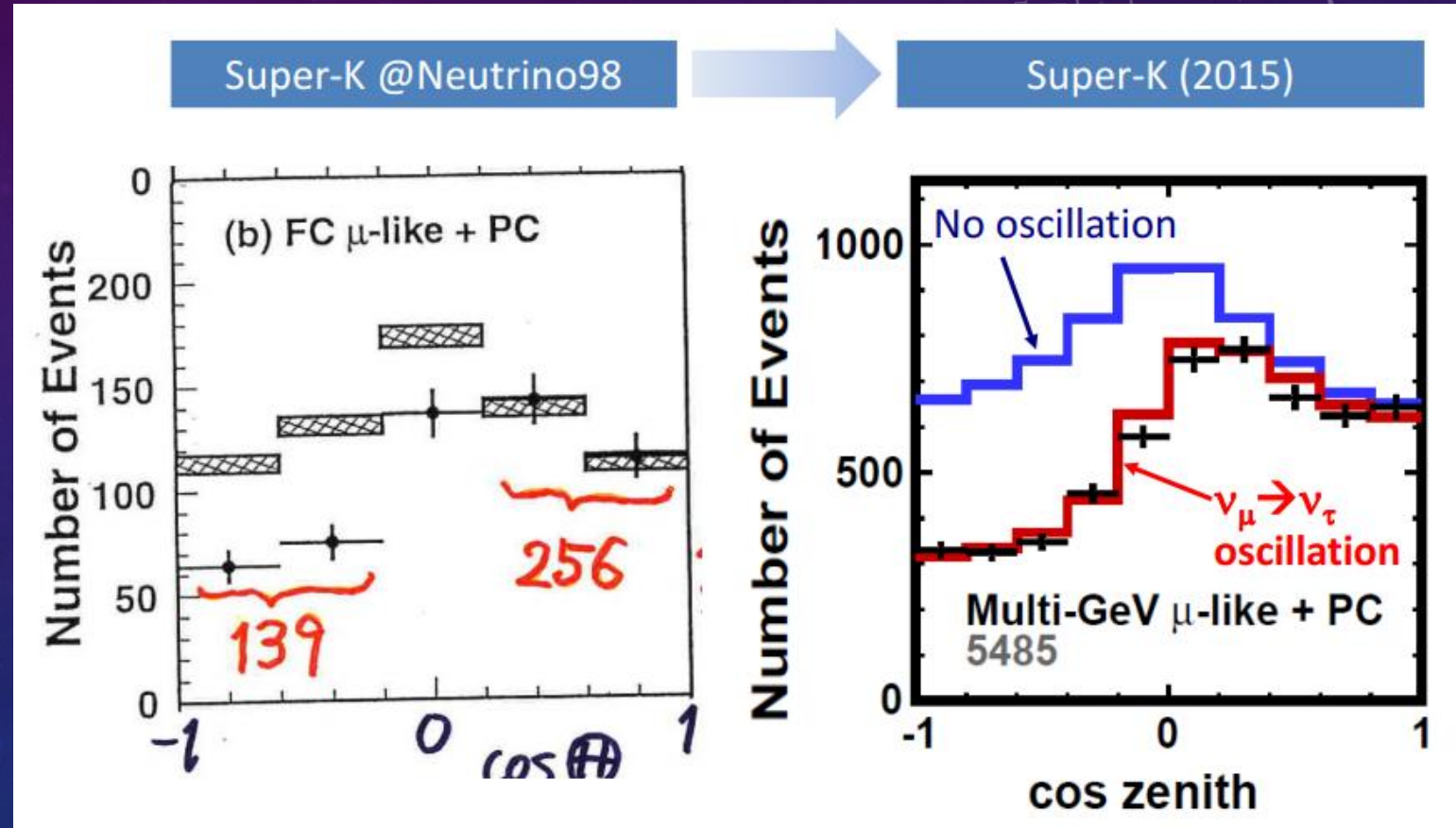
* Up/Down syst. error for μ -like

Prediction (flux calculation $\dots \lesssim 1\%$
1km rock above SK $\dots 1.5\%$) 1.8%

Data (Energy calib. for $\uparrow\downarrow \dots 0.7\%$
Non ν Background $\dots < 2\%$) 2.1%

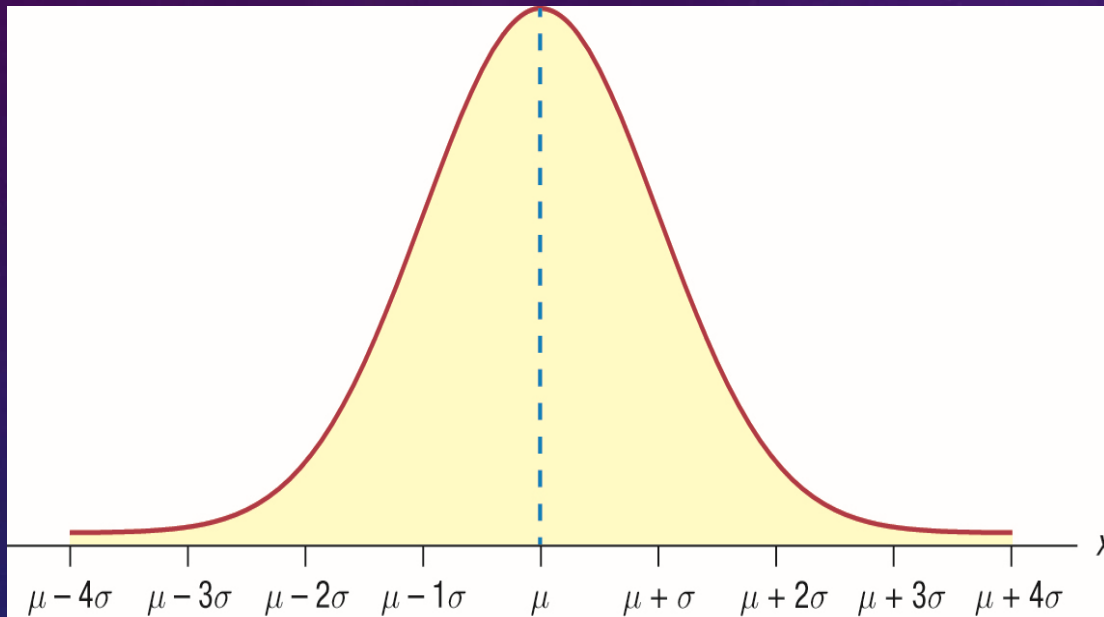
MORE DATA

- 1998 data contains 531 events
- 2015 data contains 5485 events
- Heaviest neutrino $\sim 10,000,000$ times smaller than electron
- ν_μ oscillate maximally to ν_τ



© Takaaki Kajita, Nobel Lecture 2015

GAUSSIAN DISTRIBUTION

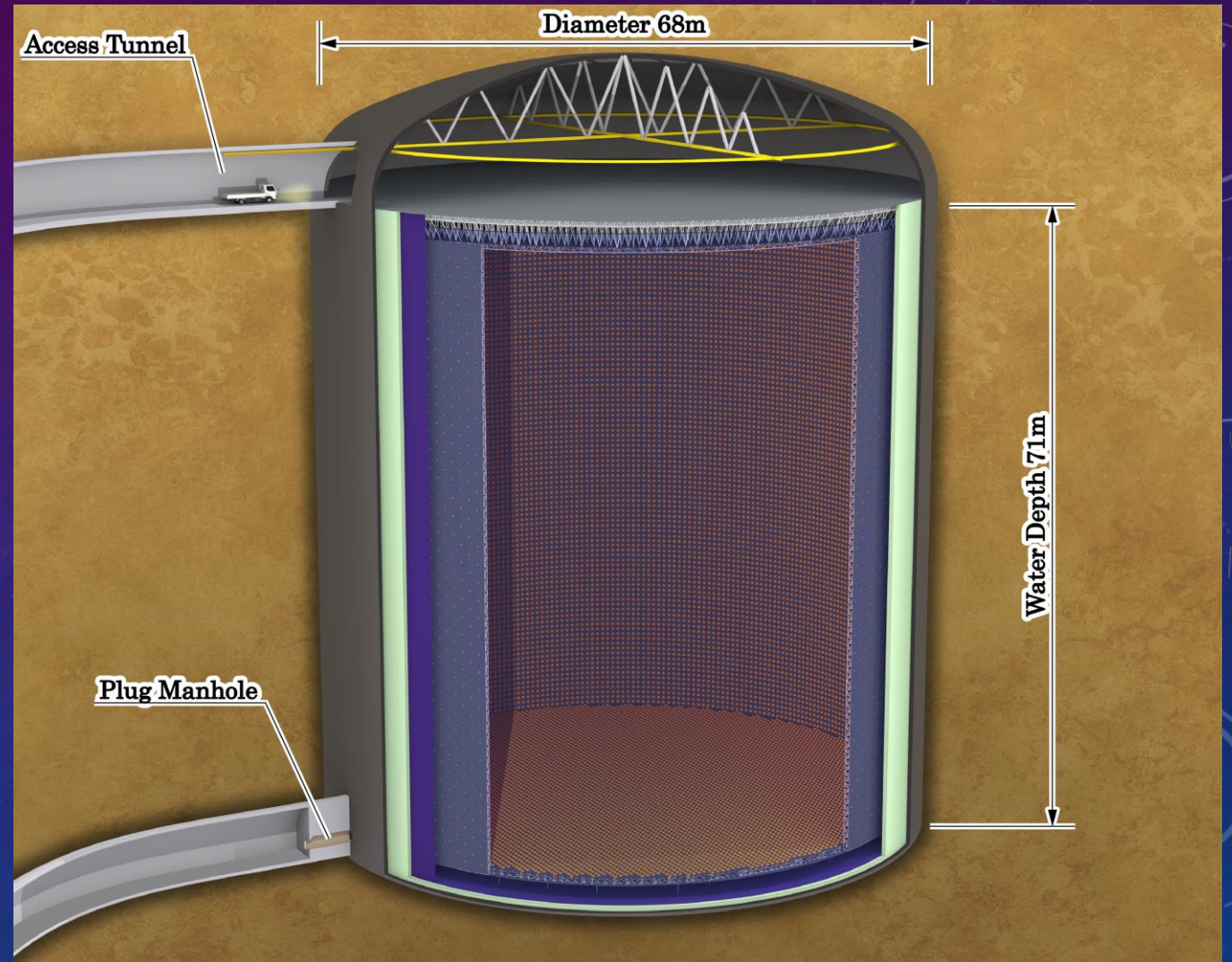


$\mu \pm N\sigma$	% within range	Daily Frequency
1	68.26894921%	Twice a week
2	95.44997361%	Every three weeks
3	99.73002039%	Yearly
4	99.99366575%	Every 43 years
5	99.99994267%	Every 4776 years
6	99.99999980%	Every 1.38 M years

© Brian Murphy & Rachelle Barr, University of Wisconsin

WHAT NEXT

- Even bigger detector, Hyper-Kamiokande
 - To start taking data in second half of 2020's
 - 50,000 -> 260,000 tons of water
 - 11,000 -> 40,000 PMT
- Goals are to
 - Order neutrino masses
 - CP violation measurement
 - Proton decay



ACKNOWLEDGEMENTS

- Takaaki Kajita – Nobel Lecture. NobelPrize.org. Nobel Media AB 2021. Thu. 4 Mar 2021. <https://www.nobelprize.org/prizes/physics/2015/kajita/lecture/>
- Hyper Kamiokande Goals – <http://www.hyper-k.org/en/physics.html>
- Neutrino Oscillation Calculations – <http://www.ps.uci.edu/~superk/oscmath2.html>
- Kamiokande Expectations – K. Hirata et al, Phys.Lett.B 205 (1988) 416

THREE NEUTRINO OSCILLATIONS

