DISCOVERY OF NEUTRINO OSCILLATIONS

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KAMIOKANDE

- Kamioka Nucleon Decay Experiment
 - 3000 tons of water
 - 1000 50cm PMTs
 - 1000m underground
- Built in 1983 to detect proton decay
- Cherenkov radiation detector

- By 1986, no evidence of proton decay
- Suspected that software was not good enough
- New software was tested on single Cherenkovring events
- Unexpected result
 - Much fewer v_{μ} than expected
 - Audit of software showed it was right

COSMIC RAY DECAY

- Cosmic Ray enters the atmosphere
- Produces a $\pi^+ \rightarrow \mu^+ + \nu_{\mu}$
- $\mu^+ \rightarrow e^+ + \nu_e + \overline{\nu}_{\mu}$
- Should expect ~2 v_{μ} per v_{e}



RESULTS OF KAMIOKANDE

K. Hirata et al, Phys.Lett.B 205 (1988) 416.		
	Data	Prediction
v _e events	93	88.5
v_{μ} events	85	144.0

WHY DOES NEUTRINO OSCILLATION IMPLY MASS

From the time - dependent Schrödinger equation :

$$\begin{pmatrix} \nu_{\mathrm{l}}(\vec{x},t) \\ \nu_{2}(\vec{x},t) \end{pmatrix} = e^{i\vec{p}\cdot\vec{x}} \begin{pmatrix} e^{-iE_{\mathrm{l}}t} | \nu_{\mathrm{l}}(0) \\ e^{-iE_{2}t} | \nu_{2}(0) \rangle \end{pmatrix}$$
$$= e^{i\vec{p}\cdot\vec{x}} \begin{pmatrix} e^{-iE_{\mathrm{l}}t} & 0 \\ 0 & e^{-iE_{2}t} \end{pmatrix} \begin{pmatrix} | \nu_{\mathrm{l}}(0) \\ \nu_{2}(0) \rangle \end{pmatrix}$$

Using the relation between mass and flavor eigenstates : $\begin{pmatrix} |\nu_{\mu}(\vec{x},t)\rangle \\ |\nu_{-}(\vec{x},t)\rangle \end{pmatrix} = e^{i\vec{p}\cdot\vec{x}} \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} e^{-iE_{1}t} & 0 \\ 0 & e^{-iE_{2}t} \end{pmatrix} \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} |\nu_{\mu}(0)\rangle \\ |\nu_{\tau}(0)\rangle \end{pmatrix}$

If
$$|\nu_{\mu}(0)\rangle = 1$$
 and $|\nu_{\tau}(0)\rangle = 0$:
 $||\nu_{\tau}(\vec{x},t)\rangle|^{2} = \sin^{2}(2\theta)\sin^{2}\frac{(E_{2}-E_{1})t}{2} \equiv P(\nu_{\mu} \rightarrow \nu_{\tau})$

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WHY DOES NEUTRINO OSCILLATION IMPLY MASS

If $E_1, E_2 >> m_1, m_2$:

$$E_2 - E_1 = \sqrt{m_2^2 + p^2} - \sqrt{m_1^2 + p^2} \approx \frac{m_2^2 - m_1^2}{2p}$$

and

$$t \approx |\vec{x}| \equiv L,$$

$$p \approx E$$

Therefore :

$$P(\nu_{\mu} \to \nu_{\tau}) \approx \sin^2(2\theta) \sin^2\left(\frac{\Delta m^2}{4}\frac{L}{E}\right)$$

- So, the probability of a neutrino oscillating is a function of: θ , L, E, and Δm
- If Δm = 0 neutrinos will have a zero probability to oscillate

NEUTRINO OSCILLATION CONDITIONS

- In the case of $v_{\mu} \rightarrow v_{\tau}$ oscillation
 - Blue curve represents probability to remain v_{μ}
 - Red curve represents probability to become v_{τ}
- At short distances L, neutrinos will have a very low probability to oscillate



NEUTRINO OSCILLATION CONDITIONS



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- Should observe a deficit in v_{μ} that pass through the earth
 - Upward going v_{μ}
- Want to create conditions to observe oscillations
- Kamiokande was not enough to be conclusive
- Need a much larger detector

SUPER-KAMIOKANDE



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- 3000 -> 50,000 ton water Cherenkov Detector
- 1000 -> 13,000 PMT
- 1000m underground

POSSIBLE NEUTRINO EVENTS



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POSSIBLE NEUTRINO EVENTS



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RESULTS

- $\cos\theta = 1$ down-going neutrino
- Shaded boxes are predictions
- Crosses are observations
- All things considered 6.2 o
 - ~1 in 1.8 billion



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MORE DATA

- 1998 data contains 531 events
- 2015 data contains 5485 events
- Heaviest neutrino ~10,000,000 times smaller than electron
- v_{μ} oscillate maximally to v_{τ}



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WHAT NEXT

- Even bigger detector, Hyper-Kamiokande
 - To start taking data in second half of 2020's
 - 50,000 -> 260,000 tons of water
 - 11,000 -> 40,000 PMT
 - Intends to order neutrino masses



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WHY THIS TOPIC?

QUESTIONS?